

2015

MATHEMATICS

( Major )

Paper : 5.3

( Spherical Trigonometry and Astronomy .)

Full Marks : 60

Time : 3 hours

*The figures in the margin indicate full marks  
for the questions*

1. Answer *all* the questions : 1×7=7
- (a) Define celestial horizon of celestial sphere and name the poles of it.
  - (b) What do you mean by great circle and what is its connection in determining celestial coordinates?
  - (c) Before postulating the Kepler's planetary laws, who was the immediate predecessor of Kepler?
  - (d) Mention a similar property of a small circle and a great circle for a sphere.

- (e) Define the polar triangle of a spherical triangle.
- (f) What is the difference in totality of lunar and solar eclipses?
- (g) State the Cassini's hypothesis of refraction.

2. Answer the following questions : 2×4=8

- (a) Mention the physical situations when a lunar eclipse can occur.
- (b) Show that in a spherical triangle
$$\pi < A + B + C < 3\pi$$
- (c) Show that the zenith distance of a star is the compliment of the altitude.
- (d) Defining sidereal time and solar time, distinguish them clearly.

3. Answer any *three* of the following : 5×3=15

- (a) If  $V_1$  and  $V_2$  are the linear velocities of a planet at perihelion and aphelion respectively, prove that

$$V_1 : V_2 = (1 + e) : (1 - e)$$

(b) If in a spherical triangle  $ABC$ ,  $A = a$ , show that  $B$  and  $b$  are equal or supplemental.

(c) Show that the right ascension (RA)  $\alpha$  and declination  $\delta$  of the sun is connected by the relation

$$\tan \delta = \tan \epsilon \sin \alpha$$

(d) Prove that if at a certain instant the declination of a star is unaffected by refraction, the azimuth of the star is then maximum and the star culminates between the pole and the zenith.

4. In any spherical triangle, prove that

$$\frac{\sin A}{\sin a} = \frac{\sin B}{\sin b} = \frac{\sin C}{\sin c}$$

Hence prove that

$$\frac{\sin(A+B)}{\sin C} = \frac{\cos a + \cos b}{1 + \cos c} \quad 6+4=10$$

5. Define true, mean and eccentric anomalies for planet's motion. Deduce the Kepler's equation  $m = \phi - e \sin \phi$ , where the symbols have their usual meanings. 3+7=10

Or

Describe the situation when a solar eclipse can occur. Prove that at the instant of conjunction in right ascension, the ratio of the distances of the sun from the moon and the earth is

$$\frac{\{\sin P_1 - \sin P \cos(\delta - \delta_1)\}}{\sin P_1}$$

where  $\delta$  and  $\delta_1$  are the declinations of the sun and the moon,  $P$  and  $P_1$  are their horizontal parallaxes (the square of  $\sin P$  is neglected).

3+7=10

6. (a) Show that the refraction of the zenith distance of a star is  $k \tan z$ , where  $z$  is the apparent zenith distance and  $k$  is a constant. 5

- (b) Show that the parallax in declination of a planet observed from a place in latitude  $\phi$  vanishes if  $\tan \phi = \cos H \tan \delta$ ,  $\delta$  and  $H$  being the planet's declination and hour-angle respectively, and the earth being assumed spherical. 5

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